# Chapter III: A cryptic Mesoarchean terrane in the basement to the Central African Copperbelt: U-Pb isotope evidence from detrital and xenocrystic zircons in the Muva and Katangan sequences* 

C. RAINAUD ${ }^{1}$, S. MASTER ${ }^{1}$, R.A. ARMSTRONG ${ }^{2}$ \& L.J. ROBB ${ }^{1}$<br>${ }^{1}$ Economic Geology Research Institute/Hugh Allsopp Laboratory, School of Geosciences, University of the Witwatersrand, Pvt. Bag 3, WITS 2050, Johannesburg, South Africa.<br>${ }^{2}$ Research School of Earth Sciences, The Australian National University, Canberra, ACT 0200, Australia.


#### Abstract

In a study of the geochronology of the Katangan Sequence and its basement in the Central African Copperbelt (Rainaud et al., 1999), detrital and xenocrystic zircons from Muva quartzites and Katangan lapilli tuffs, were dated using the SHRIMP. A detrital population (dated between 3007 and 3031 Ma ) and a group of xenocrystic zircons aged between 3169 and 3225 Ma provide the first evidence for the existence of a Mesoarchean basement beneath the Central African Copperbelt.


## 1. Introduction

The Central African Copperbelt, hosted by the Katangan Sequence, is situated in Zambia and the Katanga Province of the Democratic Republic of Congo (D.R.C.). The Katangan Sequence, which is subdivided into the Roan, Lower and Upper Kundelungu Supergroups, consists mainly of metasediments with minor mafic tuffs and sills (Figure 1). Its deposition took

[^0]place between 880 and 620 Ma (Armstrong et al., 1999; Cahen et al., 1984). The exposed basement to the Katangan Sequence consists of a Paleoproterozoic magmatic arc terrain dated at between 2050 Ma and 1800 Ma (Rainaud et al., 1999). On this basement the (as yet undated) Muva supracrustal succession of conglomerates, orthoquartzites and shales was deposited. This basement was then intruded by the 880 Ma Nchanga Granite, followed shortly by the deposition of the Katangan Sequence (Armstrong et al., 1999). To the west, the Katangan Sequence is flanked by the c. 1300-1000 Ma Kibaran Belt, which separates it from the Neoarchean rocks of the Kasai-Congo Craton (Cahen et al., 1984; Delhal, 1991; Tack et al., 1999).


Figure 1. Simplified geological map of the eastern part of the Central African Copperbelt. After François, 1974.

## 2. Sampling

Detrital zircons were separated from a sample of crossbedded quartzite (MVQ1) from the Muva Supergroup, which was collected south of Mufulira (Zambia) at $21^{\circ} 12^{\prime} E, 12^{\circ} 36^{\prime}$ S. Numerous xenocrystic zircons were found in a lapilli tuff (S11) from the Mwashya Group in the upper part of the Roan Supergroup. This tuff was sampled at Shituru Mine ( $26^{\circ} 50^{\prime} \mathrm{E}, 1^{\circ} 01^{\prime} \mathrm{S}$ ), near Likasi, in the central part of the Lufilian Arc (D. R. C.).

## 3. Analytical techniques

U-Pb analyses were performed on the SHRIMP I and II ion microprobes at The Australian National University, Canberra. The separation of zircons was carried out at the Hugh Allsopp Laboratory, Johannesburg, using conventional techniques. The SHRIMP analytical procedure used in this study is similar to that described by Claué-Long et al., (1995). Age calculations and plotting were done using Isoplot/Ex (Ludwig, 2000). Zircons from sample MVQ1 were randomly selected for analysis; in sample S11, more than $80 \%$ of all available zircons were analysed. In the following age interpretations only isotopic ratios that are $10 \%$ or less discordant were considered as reliable age indicators.

## 4. Detrital zircons

52 U-Pb analyses were carried out on 49 detrital zircons from sample MVQ1. Of these analyses, 49 were $10 \%$ or less discordant in terms of ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ages. The results are plotted on a concordia diagram in Figure 2, and the age distributions are shown on a histogram plot as an inset. The detrital zircons form several distinct populations, which range in age from 3180 to 1941 Ma . The youngest detrital zircons (22\% of the population) form a cluster of ages peaking at 1990 Ma , but which range from $2099 \pm 15 \mathrm{Ma}$ to
$1941 \pm 40 \mathrm{Ma}$ (which is the maximum age for the Muva quartzite) (Table 1). A second cluster of ages (39\%) has a peak at about 2190 Ma , and ranges from $2297 \pm 20$ to $2114 \pm 39 \mathrm{Ma}$. A third group of ages ( $6 \%$ ) ranges from $2400 \pm 19$ to $2371 \pm 17 \mathrm{Ma}$. A fourth group of zircons ( $23 \%$ ) has ages which range from $2708 \pm 18$ to $2463 \pm 25 \mathrm{Ma}$. This group has a bimodal distribution, with peaks at around 2500 Ma and 2700 Ma . There is a last group of zircons ( $8 \%$ ) whose ages range from $3031 \pm 6$ to $3007 \pm 15 \mathrm{Ma}$, with a peak at around 3020 Ma . Finally, the oldest detrital zircon is dated at $3180 \pm 12 \mathrm{Ma}$.


Figure 2. a) ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ vs ${ }^{207} \mathrm{~Pb} /{ }^{235} / \mathrm{U}$ concordia plot of ages (Ma) of detrital zircons from the Muva quartzite, sample MVQ1. b) Histogram plot showing the relative distribution of ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ages of the detrital zircons.

## 5. Xenocrystic zircons

Numerous zircons were found in a lapilli tuff (S11) from the Mwashya Group in the Katangan Sequence. Out of the 48 zircons analysed, 43 yielded ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ages that were $10 \%$ or less discordant (Figure 3). With the exception of two zircons which yielded very discordant ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ apparent ages of $609 \pm 288 \mathrm{Ma}$ and $681 \pm 67 \mathrm{Ma}$ (Table 2), all other zircons yielded ages greater than 880 Ma , the maximum age of the Katangan Sequence (Armstrong et al., 1999). Consequently the bulk of the zircon population from
Table 1. Summary of SHRIMP U-Th-Pb zircon results for sample MVQ1

|  |  |  |  |  | Radiogenic Ratios |  |  |  |  |  | Ages (in Ma) |  |  |  |  |  | Conc. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Grain. U Th | Th/U | Pb* | ${ }^{204} \mathrm{~Pb}$ / | $\mathrm{f}_{206}$ | ${ }^{206} \mathrm{~Pb} /$ |  | ${ }^{207} \mathrm{~Pb} /$ |  | ${ }^{207} \mathrm{~Pb}$ / |  | ${ }^{206} \mathrm{~Pb} /$ |  | ${ }^{207} \mathrm{~Pb} /$ |  | ${ }^{207} \mathrm{~Pb} /$ |  |  |
| spot (ppm) (ppm) |  | ( p pm ) | ${ }^{206} \mathrm{~Pb}$ | \% | ${ }^{238} \mathrm{U}$ | $\pm$ | ${ }^{235} \mathrm{U}$ | $\pm$ | ${ }^{206} \mathrm{~Pb}$ | $\pm$ | ${ }^{238} \mathrm{U}$ | $\pm$ | ${ }^{235} \mathrm{U}$ | $\pm$ | ${ }^{206} \mathrm{~Pb}$ | $\pm$ | \% |

[^1]Table 1. Summary of SHRIMP U-Th-Pb zircon results for sample MVQ1

|  |  |  |  |  |  |  | Radiogenic Ratios |  |  |  |  |  | Ages (in Ma) |  |  |  |  |  | $\begin{gathered} \text { Conc. } \\ \% \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Grain spot | $\underset{\mathrm{ppm}}{\mathrm{u}}$ | Th ppm | Th/U | $\begin{aligned} & \mathbf{P b}^{*} \\ & \mathrm{ppm} \end{aligned}$ | ${ }^{204} \mathrm{~Pb} /$ <br> ${ }^{206} \mathrm{~Pb}$ | $\mathbf{f}_{206}$ | $\begin{aligned} & { }^{206} \mathrm{~Pb} / \\ & { }^{238} \mathrm{U} \end{aligned}$ | $\pm$ | $\begin{gathered} { }^{207} \mathrm{~Pb} / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{~Pb} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\pm$ | $\begin{aligned} & -\overline{{ }^{206} \mathrm{~Pb} /} \\ & { }^{238} \mathrm{U} \end{aligned}$ | $\pm$ | $\begin{gathered} { }^{207} \mathrm{~Pb} / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{P} \mathrm{~b} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\pm$ |  |
| 29.1 | 73 | 29 | 0. | 53 | 0.0002 | 0.317 | 0.6179 | 0.0229 | 21.235 | 0.825 | 0.2493 | 0.0020 | 3102 | 92 | 3149 | 38 | 3180 | 12 | 98 |
| 30.1 | 147 | 76 | 0.52 | 87 | 0.00025 | 0.382 | 0.5180 | 0.0147 | 13.291 | 0.420 | 0.1861 | 0.0020 | 2691 | 63 | 2701 | 30 | 2708 | 18 | 99 |
| 30.2 | 25 | 23 | 0.90 | 16 | 0.00215 | 3.287 | 0.5144 | 0.0226 | 12.559 | 0.825 | 0.1771 | 0.0077 | 2675 | 97 | 2647 | 64 | 2626 | 75 | 102 |
| 31.1 | 190 | 74 | 0.39 | 87 | 0.00015 | 0.234 | 0.4218 | 0.0105 | 8.309 | 0.231 | 0.1429 | 0.0014 | 2269 | 48 | 2265 | 25 | 2262 | 17 | 100 |
| 32.1 | 208 | 353 | 1.70 | 87 | 0.00053 | 0.812 | 0.3613 | 0.0107 | 6.117 | 0.230 | 0.1228 | 0.0024 | 1988 | 51 | 1993 | 33 | 1997 | 36 | 100 |
| 33.1 | 124 | 104 | 0.84 | 51 | 0.00032 | 0.490 | 0.3594 | 0.0105 | 5.967 | 0.206 | 0.1204 | 0.0018 | 1979 | 50 | 1971 | 30 | 196 | 27 | 10 |
| 34. | 20 | 27 | 0 | 92 | 0.00012 | 0.180 | 0.4115 | 0.0109 | 7.707 | 0.222 | 0.1358 | 0.0011 | 222 | 50 | 2197 | 26 | 217 | 15 | 10 |
| 35.1 | 166 | 56 | 0.34 | 71 | 0.00019 | 0.297 | 0.4027 | 0.0120 | 7.616 | 0.247 | 0.1372 | 0.0013 | 2181 | 55 | 2187 | 30 | 2192 | 17 | 100 |
| 36.1 | 348 | 146 | 0.42 | 138 | 0.00013 | 0.195 | 0.3720 | 0.0082 | 6.544 | 0.158 | 0.1276 | 0.0010 | 2039 | 38 | 2052 | 21 | 2065 | 13 | 99 |
| 37.1 | 110 | 124 | 1.13 | 58 | 0.00037 | 0.570 | 0.4224 | 0.0110 | 8.492 | 0.253 | 0.1458 | 0.0017 | 2271 | 50 | 2285 | 27 | 2297 | 20 | 99 |
| 38.1 | 331 | 157 | 0.47 | 128 | 0.00017 | 0.268 | 0.3665 | 0.0096 | 6.683 | 0.191 | 0.1322 | 0.0012 | 2013 | 45 | 2070 | 26 | 2128 | 16 | 95 |
| 39.1 | 75 | 128 | 71 | 66 | 0.00035 | 0.532 | 0.6118 | 0.0246 | 18.867 | 0.802 | 0.2237 | 0.0021 | 3077 | 99 | 3035 | 42 | 3007 | 15 | 102 |
| 39. | 33 | 43 | 1.33 | 28 | 0.00011 | 0.166 | 0.6288 | 0.0284 | 19.535 | 1.002 | 0.2253 | 0.0044 | 3145 | 113 | 3069 | 51 | 3019 | 32 | 104 |
| 40.1 | 285 | 1298 | 4.56 | 70 | 0.00354 | 5.423 | 0.1970 | 0.0050 | 3.454 | 0.194 | 0.1272 | 0.0060 | 1159 | 27 | 1517 | 45 | 2059 | 85 | 56 |
| 41.1 | 266 | 463 | 1.74 | 98 | 0.00048 | 0.734 | 0.3304 | 0.0079 | 5.639 | 0.170 | 0.1238 | 0.0019 | 1841 | 38 | 1922 | 26 | 2011 | 28 | 92 |
| 42.1 | 259 | 142 | 0.55 | 157 | 0.00022 | 0.331 | 0.5309 | 0.0126 | 13.592 | 0.351 | 0.1857 | 0.0014 | 2745 | 53 | 2722 | 25 | 2704 | 13 | 102 |
| 43.1 | 187 | 105 | 0.56 | 90 | 0.00014 | 0.260 | 0.4314 | 0.0204 | 8.172 | 0.418 | 0.1374 | 0.0020 | 2312 | 92 | 2250 | 47 | 2195 | 26 | 105 |
| 44.1 | 307 | 189 | 0.62 | 100 | 0.00105 | 1.971 | 0.3091 | 0.0140 | 5.624 | 0.301 | 0.1319 | 0.0031 | 1737 | 69 | 1920 | 47 | 2124 | 42 | 82 |
| 45.1 | 221 | 129 | 0.58 | 90 | 0.00029 | 0.541 | 0.3673 | 0.0167 | 6.059 | 0.301 | 0.1197 | 0.0018 | 2017 | 79 | 1984 | 44 | 1951 | 27 | 103 |
| 46.1 | 144 | 168 | 1.17 | 98 | 0.00007 | 0.138 | 0.5327 | 0.0244 | 13.500 | 0.653 | 0.1838 | 0.0020 | 2753 | 103 | 2715 | 47 | 2688 | 18 | 102 |
| 47.1 | 143 | 70 | 0.49 | 60 | 0.00012 | 0.235 | 0.3888 | 0.0187 | 6.747 | 0.348 | 0.1259 | 0.0017 | 2117 | 87 | 2079 | 47 | 2041 | 25 | 104 |
| 48.1 | 406 | 222 | 0.55 | 183 | 0.00009 | 0.162 | 0.4101 | 0.0178 | 7.735 | 0.347 | 0.1368 | 0.0010 | 2216 | 82 | 2201 | 41 | 2187 | 13 | 101 |
| 49.1 | 263 | 100 | 0.38 | 183 | 0.00007 | 0.126 | 0.6114 | 0.0274 | 19.110 | 0.877 | 0.2267 | 0.0012 | 3076 | 111 | 3047 | 45 | 3029 | 9 | 102 |
| 50.1 | 154 | 115 | 0.74 | 85 | 0.00004 | 0.066 | 0.4718 | 0.0229 | 10.069 | 0.515 | 0.1548 | 0.0017 | 2491 | 101 | 2441 | 48 | 2400 | 19 | 104 |

[^2]tuff sample is interpreted to be xenocrystic in origin. These zircons form several distinct age populations ranging from $1018 \pm 27 \mathrm{Ma}$ to $3225 \pm 11 \mathrm{Ma}$. Some of these zircon populations have ages that overlap those from the detrital zircon population in the Muva quartzite. However, some groups from the detrital suite are not represented in the suite and vice versa. The xenocrystic suite include a population of Mesoproterozoic zircons which is completely absent from the detrital zircon population in the Muva quartzites. The youngest xenocrystic zircon population, comprising 5 zircons (12\% of the 43 analyses used), is dated at between $1537 \pm 89$ and $1018 \pm 27 \mathrm{Ma}$. A second group of 22 zircons (51\%) has Paleoproterozoic ages between $2105 \pm 25$ and $1791 \pm 21 \mathrm{Ma}$, with a peak at c. 1860 Ma . One zircon grain provides an age of $2624 \pm 9 \mathrm{Ma}$, while another is dated at $3021 \pm 34 \mathrm{Ma}$. Finally, there is a large group of 14 zircons ( $33 \%$ ) which are dated at between $3225 \pm 11$ and $3169 \pm 13$ Ma.


Figure 3. a) ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ vs ${ }^{207} \mathrm{~Pb} /{ }^{235} / \mathrm{U}$ concordia plot of ages (Ma) of inherited xenocrystic zircons from lapilli tuff (sample S11), Mwashya Group, Katangan Sequence. b) Histogram plot showing the relative distribution of ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ages of the xenocrystic zircons.
Table 2. Summary of SHRIMP U-Th-Pb zircon results for sample S 11

|  |  |  |  |  |  |  | Radiogenic Ratios |  |  |  |  |  | Ages (in Ma) |  |  |  |  |  | Conc \% |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Grain. spot | $\begin{gathered} U \\ (\mathrm{ppm}) \end{gathered}$ | $\begin{gathered} T h \\ (\mathrm{ppm}) \end{gathered}$ | Th/U | $\begin{gathered} P b^{*} \\ (\mathrm{ppm}) \end{gathered}$ | $\begin{aligned} & { }^{204} \mathrm{~Pb} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\begin{gathered} f_{206} \\ \% \end{gathered}$ | $\begin{gathered} { }^{206} \mathrm{P} \text { b } / \\ { }^{238} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{gathered} { }^{207} \mathrm{P} \text { b } / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{~Pb} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\pm$ | $\begin{gathered} { }^{206} \mathrm{P} \text { b } / \\ { }^{238} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{gathered} { }^{207} \mathrm{P} \text { b } / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{~Pb} / \\ & { }^{206} \mathrm{P} \mathrm{~b} \end{aligned}$ | $\pm$ |  |
| 1.1 | 269 | 288 | 1.07 | 105 | 0.00010 | 0.2 | 0.3224 | 0.0056 | 5.153 | 0.100 | 0.1159 | 0.0008 | 1801 | 28 | 1845 | 17 | 1894 | 12 | 95 |
| 2.1 | 194 | 197 | 1.02 | 79 | 0.00007 | 0.1 | 0.3348 | 0.0063 | 5.230 | 0.111 | 0.1133 | 0.0008 | 1862 | 31 | 1858 | 18 | 1853 | 13 | 101 |
| 3.1 | 271 | 231 | 0.85 | 54 | 0.00010 | 0.1 | 0.1730 | 0.0034 | 1.744 | 0.044 | 0.0731 | 0.0010 | 1029 | 19 | 1025 | 16 | 1018 | 27 | 101 |
| 4.1 | 215 | 149 | 0.69 | 82 | 0.00008 | 0.1 | 0.3397 | 0.0058 | 5.315 | 0.102 | 0.1135 | 0.0008 | 1885 | 28 | 1871 | 17 | 1856 | 12 | 102 |
| 5.1 | 94 | 91 | 0.97 | 38 | 0.00003 | 0 | 0.3315 | 0.0071 | 5.116 | 0.128 | 0.1119 | 0.0012 | 1846 | 35 | 1839 | 21 | 1831 | 19 | 101 |
| 6.1 | 152 | 122 | 0.80 | 66 | 0.00002 | 0 | 0.3703 | 0.0060 | 6.215 | 0.113 | 0.1218 | 0.0008 | 2031 | 28 | 2007 | 16 | 1982 | 11 | 103 |
| 7.1 | 176 | 222 | 1.26 | 76 | 0.00006 | 0.1 | 0.3380 | 0.0069 | 5.303 | 0.122 | 0.1138 | 0.0010 | 1877 | 33 | 1869 | 20 | 1861 | 15 | 101 |
| 8.1 | 93 | 100 | 1.07 | 78 | 0.00123 | 1.9 | 0.6363 | 0.0125 | 21.715 | 0.485 | 0.2475 | 0.0021 | 3174 | 49 | 3171 | 22 | 3169 | 13 | 100 |
| 9.1 | 89 | 72 | 0.81 | 73 | 0.00016 | 0.3 | 0.6489 | 0.0142 | 22.658 | 0.532 | 0.2532 | 0.0015 | 3224 | 56 | 3212 | 23 | 3205 | 10 | 101 |
| 10.1 | 102 | 68 | 0.66 | 44 | 0.00001 | 0 | 0.3768 | 0.0077 | 6.655 | 0.158 | 0.1281 | 0.0013 | 2061 | 36 | 2067 | 21 | 2072 | 17 | 100 |
| 11.1 | 159 | 116 | 0.73 | 60 | 0.00004 | 0.1 | 0.3299 | 0.0060 | 5.084 | 0.102 | 0.1118 | 0.0008 | 1838 | 29 | 1834 | 17 | 1829 | 12 | 101 |
| 12.1 | 122 | 69 | 0.57 | 41 | 0.00005 | 0.1 | 0.3060 | 0.0064 | 4.619 | 0.115 | 0.1095 | 0.0012 | 1721 | 32 | 1753 | 21 | 1791 | 21 | 96 |
| 13.1 | 95 | 80 | 0.84 | 76 | 0.00003 | 0 | 0.6326 | 0.0128 | 21.854 | 0.478 | 0.2505 | 0.0015 | 3160 | 51 | 3177 | 21 | 3188 | 10 | 99 |
| 14.1 | 136 | 163 | 1.20 | 118 | 0.00023 | 0.4 | 0.6422 | 0.0140 | 22.312 | 0.508 | 0.2520 | 0.0012 | 3198 | 55 | 3197 | 22 | 3197 | 7 | 100 |
| 15.1 | 88 | 79 | 0.90 | 71 | 0.00007 | 0.1 | 0.6246 | 0.0132 | 21.958 | 0.517 | 0.2550 | 0.0020 | 3128 | 53 | 3182 | 23 | 3216 | 13 | 97 |
| 16.1 | 130 | 169 | 1.30 | 118 | 0.00025 | 0.4 | 0.6601 | 0.0333 | 22.805 | 1.195 | 0.2506 | 0.0022 | 3268 | 131 | 3219 | 52 | 3188 | 14 | 103 |
| 17.1 | 226 | 118 | 0.52 | 44 | 0.00021 | 0.3 | 0.1841 | 0.0090 | 1.892 | 0.112 | 0.0745 | 0.0021 | 1090 | 49 | 1078 | 40 | 1056 | 59 | 103 |
| 18.1 | 146 | 111 | 0.76 | 55 | 0.00069 | 1.1 | 0.3310 | 0.0157 | 5.075 | 0.310 | 0.1112 | 0.0037 | 1843 | 76 | 1832 | 53 | 1819 | 62 | 101 |
| 19.1 | 80 | 58 | 0.73 | 67 | 0.00001 | 0 | 0.6740 | 0.0318 | 23.092 | 1.133 | 0.2485 | 0.0022 | 3321 | 124 | 3231 | 49 | 3175 | 14 | 105 |
| 20.1 | 73 | 51 | 0.69 | 59 | 0.00020 | 0.3 | 0.6498 | 0.0312 | 22.526 | 1.121 | 0.2514 | 0.0022 | 3228 | 123 | 3207 | 50 | 3194 | 14 | 101 |
| 21.1 | 90 | 76 | 0.84 | 73 | 0.00014 | 0.2 | 0.6379 | 0.0190 | 22.104 | 0.727 | 0.2513 | 0.0027 | 3181 | 75 | 3188 | 32 | 3193 | 17 | 100 |
| 22.1 | 53 | 79 | 1.51 | 17 | 0.00025 | 0.4 | 0.2394 | 0.0093 | 3.150 | 0.201 | 0.0954 | 0.0044 | 1384 | 48 | 1445 | 50 | 1537 | 89 | 90 |
| 23.1 | 145 | 133 | 0.92 | 116 | 0.00006 | 0.1 | 0.6202 | 0.0176 | 21.399 | 0.689 | 0.2503 | 0.0030 | 3111 | 70 | 3157 | 32 | 3186 | 19 | 98 |
| 24.1 | 78 | 148 | 1.89 | 23 | 0.00053 | 0.9 | 0.2064 | 0.0069 | 2.360 | 0.143 | 0.0829 | 0.0039 | 1210 | 37 | 1231 | 44 | 1267 | 94 | 96 |
| 25.1 | 225 | 89 | 0.40 | 54 | 0.00019 | 0.3 | 0.2310 | 0.0062 | 2.649 | 0.100 | 0.0832 | 0.0020 | 1340 | 32 | 1315 | 28 | 1273 | 46 | 105 |

[^3]Table 2. Summary of SHRIMP U-Th-Pbzircon results for sample S 11

| Grain. spot | $\begin{gathered} U \\ (\mathrm{p} p \mathrm{~m}) \end{gathered}$ | $\begin{gathered} \mathrm{Th} \\ (\mathrm{ppm}) \end{gathered}$ | Th/U | $\begin{gathered} \mathrm{Pb} \mathrm{~b}^{*} \\ (\mathrm{ppm}) \end{gathered}$ | ${ }^{204} \mathrm{~Pb} /$ <br> ${ }^{206} \mathrm{P}$ b | $\begin{gathered} \mathbf{f}_{206} \\ \% \end{gathered}$ | $\begin{gathered} \hline{ }^{206} \mathrm{~Pb} / \\ { }^{238} \mathrm{U} \end{gathered}$ | Radiogenic Ratios |  |  |  | Ages (in Ma) |  |  |  |  |  |  | Conc |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{P} \text { b/ } \\ & { }^{235} \mathrm{U} \end{aligned}$ | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{~Pb} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\pm$ | $\begin{gathered} { }^{206} \mathrm{~Pb} / \\ { }^{238} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{gathered} { }^{207} \mathrm{~Pb} / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\pm$ | $\begin{aligned} & { }^{207} \mathrm{~Pb} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\pm$ |  |
| 26.1 | 230 | 34 | 0.15 | 132 | 0.00007 | 0.1 | 0.5191 | 0.0158 | 17.067 | 0.551 | 0.2385 | 0.0018 | 2695 | 67 | 2939 | 31 | 3110 | 12 | 87 |
| 27.1 | 219 | 194 | 0.88 | 95 | 0.00012 | 0.2 | 0.3683 | 0.0090 | 6.045 | 0.163 | 0.1191 | 0.0011 | 2021 | 42 | 1982 | 24 | 1942 | 16 | 104 |
| 28.1 | 96 | 72 | 0.75 | 79 | 0.00015 | 0.2 | 0.6532 | 0.0255 | 22.561 | 0.940 | 0.2505 | 0.0027 | 3241 | 100 | 3208 | 41 | 3188 | 17 | 102 |
| 29.1 | 210 | 119 | 0.57 | 92 | 0.00013 | 0.2 | 0.3924 | 0.0108 | 7.061 | 0.229 | 0.1305 | 0.0018 | 2134 | 50 | 2119 | 29 | 2105 | 25 | 101 |
| 30.1 | 359 | 222 | 0.62 | 137 | 0.00014 | 0.2 | 0.3447 | 0.0093 | 5.235 | 0.153 | 0.1101 | 0.0009 | 1909 | 45 | 1858 | 25 | 1802 | 16 | 106 |
| 31.1 | 435 | 36 | 0.08 | 212 | 0.00007 | 0.1 | 0.4725 | 0.0107 | 11.523 | 0.275 | 0.1769 | 0.0009 | 2495 | 47 | 2566 | 23 | 2624 | 9 | 95 |
| 32.2 | 178 | 82 | 0.46 | 65 | 0.00030 | 0.5 | 0.3435 | 0.0100 | 5.693 | 0.213 | 0.1202 | 0.0024 | 1904 | 48 | 1930 | 33 | 1959 | 36 | 97 |
| 33.1 | 401 | 214 | 0.53 | 213 | 0.00041 | 0.5 | 0.4556 | 0.0111 | 13.197 | 0.341 | 0.2101 | 0.0013 | 2420 | 49 | 2694 | 25 | 2906 | 10 | 83 |
| 34.1 | 266 | 115 | 0.43 | 163 | 0.00012 | 0.2 | 0.5039 | 0.0113 | 17.321 | 0.430 | 0.2493 | 0.0020 | 2631 | 49 | 2953 | 24 | 3180 | 13 | 83 |
| 35.1 | 25 | 20 | 0.79 | 10 | 0.00079 | 1.2 | 0.3509 | 0.0165 | 5.492 | 0.554 | 0.1135 | 0.0095 | 1939 | 79 | 1899 | 91 | 1856 | 159 | 104 |
| 36.1 | 207 | 318 | 1.54 | 101 | 0.00020 | 0.3 | 0.3672 | 0.0090 | 5.592 | 0.166 | 0.1105 | 0.0016 | 2016 | 42 | 1915 | 26 | 1807 | 26 | 112 |
| 37.1 | 286 | 263 | 0.92 | 43 | 0.00027 | 0.5 | 0.1291 | 0.0028 | 1.107 | 0.044 | 0.0622 | 0.0019 | 782 | 16 | 757 | 21 | 681 | 67 | 115 |
| 38.1 | 92 | 61 | 0.66 | 12 | 0.00111 | 2 | 0.1229 | 0.0037 | 1.019 | 0.132 | 0.0602 | 0.0073 | 747 | 21 | 714 | 68 | 609 | 288 | 123 |
| 38.2 | 117 | 94 | 0.81 | 17 | 0.00001 | 0 | 0.1298 | 0.0033 | 1.281 | 0.054 | 0.0716 | 0.0022 | 787 | 19 | 837 | 24 | 974 | 63 | 81 |
| 39.1 | 321 | 349 | 1.09 | 142 | 0.00019 | 0.3 | 0.3622 | 0.0072 | 5.730 | 0.136 | 0.1148 | 0.0012 | 1992 | 34 | 1936 | 21 | 1876 | 20 | 106 |
| 40.1 | 18 | 2 | 0.10 | 12 | 0.00006 | 0.1 | 0.5951 | 0.0326 | 18.506 | 1.128 | 0.2256 | 0.0047 | 3010 | 133 | 3016 | 60 | 3021 | 34 | 100 |
| 41.1 | 95 | 67 | 0.71 | 38 | 0.00008 | 0.1 | 0.3515 | 0.0111 | 5.630 | 0.203 | 0.1162 | 0.0016 | 1942 | 53 | 1921 | 32 | 1898 | 25 | 102 |
| 42.1 | 1046 | 1125 | 1.08 | 178 | 0.00130 | 2 | 0.1409 | 0.0026 | 2.075 | 0.066 | 0.1068 | 0.0026 | 850 | 14 | 1141 | 22 | 1746 | 44 | 49 |
| 42.2 | 533 | 333 | 0.62 | 155 | 0.00018 | 0.3 | 0.2586 | 0.0054 | 4.197 | 0.105 | 0.1177 | 0.0013 | 1482 | 28 | 1673 | 21 | 1922 | 21 | 77 |
| 43.1 | 86 | 63 | 0.73 | 72 | 0.00001 | 0 | 0.6753 | 0.0207 | 23.380 | 0.757 | 0.2511 | 0.0018 | 3326 | 80 | 3243 | 32 | 3192 | 12 | 104 |
| 44.1 | 182 | 181 | 0.99 | 78 | 0.00012 | 0.2 | 0.3526 | 0.0079 | 5.493 | 0.158 | 0.1130 | 0.0017 | 1947 | 38 | 1899 | 25 | 1848 | 28 | 105 |
| 45.1 | 128 | 100 | 0.78 | 52 | 0.00030 | 0.5 | 0.3572 | 0.0097 | 5.459 | 0.185 | 0.1108 | 0.0019 | 1969 | 46 | 1894 | 30 | 1813 | 32 | 109 |
| 46.1 | 219 | 459 | 2.09 | 111 | 0.00012 | 0.2 | 0.3417 | 0.0073 | 5.368 | 0.139 | 0.1140 | 0.0014 | 1895 | 35 | 1880 | 22 | 1863 | 22 | 102 |
| 47.1 | 95 | 64 | 0.67 | 77 | 0.00016 | 0.2 | 0.6585 | 0.0172 | 22.815 | 0.640 | 0.2513 | 0.0018 | 3261 | 67 | 3219 | 28 | 3193 | 12 | 102 |
| 48.1 | 95 | 61 | 0.64 | 81 | 0.00001 | 0 | 0.6918 | 0.0188 | 24.458 | 0.706 | 0.2564 | 0.0018 | 3389 | 72 | 3287 | 29 | 3225 | 11 | 105 |
| 49.1 | 265 | 121 | 0.46 | 112 | 0.00009 | 0.1 | 0.3926 | 0.0093 | 6.809 | 0.176 | 0.1258 | 0.0010 | 2135 | 43 | 2087 | 23 | 2040 | 13 | 105 |
| 50.1 | 270 | 335 | 1.24 | 119 | 0.00091 | 1.4 | 0.3487 | 0.0069 | 5.471 | 0.144 | 0.1138 | 0.0017 | 1929 | 33 | 1896 | 23 | 1861 | 27 | 104 |

[^4]
## 6. Provenance of zircons

The youngest xenocrystic zircons in the Katangan tuff (1018 $\pm 27$ to $1537 \pm 89 \mathrm{Ma}$ ) span the age of the Kibaran granites ( 1375 to 1000 Ma ; Tack et al., 1999), and indicates the presence of Kibaran magmatic rocks beneath the central part of the Lufilian Arc. The absence of Kibaran-aged zircons from the detrital population in the Muva quartzite (which is derived from a much wider area than that sampled by the tuff) indicates that the Muva quartzites were most probably deposited before $1537 \pm 89 \mathrm{Ma}$. The large population of Paleoproterozoic detrital and xenocrystic zircons, dated between 1791 and 2105 Ma , overlaps the time period ( 2050 to 1800 Ma ) of the Ubendian magmatic arc terrain that constitutes the Bangweulu Block and the exposed basement in the Zambian Copperbelt (Cahen et al., 1984; Rainaud et al., 1999). A younger group of c. 1860 Ma xenocrystic zircons from the tuff is not represented in the detrital zircon population from the Muva quartzite. The 2297 to 2114 Ma detrital zircons (which are also absent from the xenocrystic suite) are from an unknown source, since there are no dated rocks of this age in the immediate vicinity. They may be derived from the Magondi Belt of Zimbabwe, which has been dated at between $2160 \pm 100$ and $2120 \pm 40 \mathrm{Ma}$ (Höhndorf and Vetter, 1999; Master, 1991; Schidlowski and Todt, 1998). The earliest Proterozoic suite of detrital zircons, dated at between 2400 and 2371 Ma , may have been derived from the Luiza metasediments in the Kasai region of the Congo, which have been dated at c. 2400 Ma (Cahen et al., 1984), or from the c. 2400 Ma granulites of the Kasai-Lomami complex (Delhal et al., 1986). The largely Neoarchean suite of detrital zircons, dated between 2710 to 2460 Ma , appears to have been derived mainly from the Kasai Craton in Congo, NE Angola and NW Zambia, where granites and migmatites have been dated at 2560 to 2540 Ma (Key and Armstrong, 2000), and where 2870 Ma leucogranites were overprinted at 2600 to 2007 Ma (Delhal, 1991). Neoarchaean rocks do not appear to be abundant in beneath the Lufilian Arc, since only one xenocrystic zircon of this age ( $2624 \pm 9 \mathrm{Ma}$ ) was found in the Katangan tuff. The most enigmatic zircons from both the
detrital and xenocrystic suites are the $>3000 \mathrm{Ma}$ (Mesoarchean) grains which fall in two clusters at c. 3020 and 3200 Ma . In the whole of Central Africa, there are no rocks that have been dated at 3200 Ma , while only a few dates of c. 3000 Ma are known in widely separated regions such as Gabon (CahenVachette et al., 1988), Zimbabwe (Cahen et al., 1984) and the northern Congo-Kasai Craton (Lavreau and Deblond, 2000).

## 7. Cryptic mesoarchaean terrane

There are no exposed rocks in the immediate vicinity of the Muva quartzites (or their proximal source regions) which are between 3300 and 3000 Ma in age. In eastern Zambia, Liyungu and Vinyu (1996) have obtained Pb model ages of $3047 \pm 130 \mathrm{Ma}$ for zircons from the Chipata granite, and $2985 \pm 14 \mathrm{Ma}$ for zircons from the Lutembwe granulite. The oldest dated rocks of the Kasai sector of the Archean Congo-Kasai craton of D.R.C. and NE Angola are c. 2900 Ma (Delhal, 1991). The greenstone belts and granites of Gabon, which constitute the western part of the Congo-Kasai craton, have been dated at 3100-2900 Ma (Cahen-Vachette et al., 1988). The Zimbabwe Archean craton consists mainly of 2900 to 2600 Ma granite-greenstone terranes, with only the southernmost Tokwe terrane containing older rocks dated at between 3500 and 2950 Ma (Kusky, 1998). The Archean Tanzanian craton is dated at between 2930 and 2530 Ma (Pinna et al., 1999). The detrital zircon population in the Muva quartzite has just one older 3180 Ma Mesoarchean zircon, and several zircons ranging from 3031 to 3007 Ma. By contrast, the Katangan tuff contains abundant older Mesoarchean xenocrystic zircons dated between 3225 and 3169 Ma , and only one younger zircon dated at 3021 Ma .


Figure 4. Sketch map showing the proposed extent of the Mesoarchean Likasi terrane (heavy dashes) beneath the Central African Copperbelt, relative to some of the important tectono-stratigraphic units of the region; i.e. the Lufilian Arc (light dashes extending from Mwinilunga [Mw] through Likasi [Li] to Mufulira [Mu]), the Neoarchean Congo-Kasai Craton, the Mesoproterozoic Kibaran and Irumide orogenic belts, the Paleoproterozoic Ubendian belt and Bangweulu block [BB], and the Muva Supergroup (Mp). The cluster of diamondiferous kimberlite pipes on the Kundelungu plateau is shown as KP .

The xenocrystic zircons originated from either the partially molten source region or the wallrocks in the path of ascent of the magmas that gave rise to the sampled Katangan tuffs. The lapilli tuffs at Shituru (which are interbedded with agglomerates) are the thickest and most proximal of all the tuffs in the Mwashya Group of the Katangan Sequence (Lefebvre, 1975). Therefore the xenocrystic zircons in these tuffs represent a sample of the crust beneath the central part of the Lufilian Arc, which is buried under the tectonically thickened Katangan Sequence. The abundance of c. 3200 Ma xenocrystic zircons in this tuff ( $32 \%$ of the total population) indicates that a part of the crust beneath the central Lufilian Arc is a c. 3200 Ma terrane that we propose to call the Likasi Terrane. The almost total absence of ages ranging from 2700 to 2500 Ma in the xenocrystic zircon population (which are abundant in the detrital zircon population) implies a lack of Neoarchean crust in the Likasi Terrane. The bulk of the remaining crust is of Palaeoproterozoic (Ubendian) age, between 2100 and 1800 Ma . There is also evidence from these
xenocrystic zircons of Kibaran-aged crust ( 1300 to 1000 Ma ) in this region. Because only a single c. 3200 Ma detrital zircon was found in the Muva quartzite, it appears that the 3200 Ma crust might have been poorly exposed at the surface, which was dominated by the c. 2000 Ma Ubendian crust. The 3200 Ma crust may have been more abundant at mid- or deep crustal levels beneath the Lufilian Arc. The xenocrystic and detrital zircons from the Katangan tuffs and Muva quartzites provide the only direct evidence for the existence of this cryptic Mesoarchean crust beneath the Katangan Sequence. The occurrence of diamonds in the kimberlites of the Kundelungu Plateau to the north of the Lufilian Arc (Demaiffe et al., 1991) may be an indirect indication of the presence of an Archean crust beneath the Katangan of Central Africa. In addition to mantle xenoliths, the Kundelungu kimberlites also contain undated crustal gneiss and mica-schist xenoliths, some of which may be samples of the cryptic Likasi Terrane. In the northwestern corner of Zambia, near Mwinilunga, close to the borders with Angola and D.R.C., Neoarchean foliated granites dated at $2538 \pm 10 \mathrm{Ma}$ contain xenocrystic inherited zircons which give a mixture of ages as old as 3154 Ma (Key and Armstrong, 2000). This indicates that the c. 3200 Ma Likasi Terrane may have extended towards the southwest from the Likasi area to the Mwinilungu area, over a distance of about 300 km (Figure 4). If the 3154 Ma xenocrystic zircons of Mwinilunga are derived from the Likasi Terrane, it implies that the latter was an integral part of the Kasai-Congo craton by the latest Archean. It is further suggested that the Likasi Terrane was accreted onto the KasaiCongo craton before 2538 Ma , and that this collisional event may have been responsible for some of the granulite-facies metamorphism in the southeastern Kasai-Congo craton.

## 8. References

Armstrong, R. A., Robb, L. J., Master, S., Kruger, F. J. and Mumba, P. A. C. C., 1999, New U-Pb age constraints on the Katangan Sequence, Central African Copperbelt: Journal of African Earth Sciences, v. 28(4A), p. 6-7.

Caen-Vachette, M., Vialette, Y., Bassot, J. -P. and Vidal, P., 1988, Apport de la géochronologie isotopique à la connaissance de la géologie gabonaise: Chronique de la recherche minière, v. 491, p. 35-54.

Cahen, L., Snelling., N. J., Delhal, J., Vail, J. R., Bonhomme, M. and Ledent, D., 1984, The Geochronology and Evolution of Africa: Oxford, 512 p.

Claué-Long, J. C., Compston, W., Roberts, J. and Fanning, C. M., 1995, Two carboniferous ages: a comparison of SHRIMP zircon dating with conventional zircon ages and $40 \mathrm{Ar} / 39 \mathrm{Ar}$ analysis: Geochronology Time Scales and Global Stratigraphic Correlation, SEPM Special Publication No 54, p. 1-21.

Delhal, J., Deutsch, S. and Denoiseux, B., 1986, A Sm-Nd isotopic study of heterogeneous granulites from the Archean Kasai-Lomami gabbro-norite and charnockite complex (Zaire, Africa): Chemical Geology, v. 57, p. 235-245.

Delhal, J., 1991, Situation géochronologique 1990 du Précambrien du SudKasai et de l'Ouest-Shaba.: Tervuren, Musée royal d'Afrique centrale, Annual report 1990, p 119-125.

Demaiffe, D., Fieremans, M. and Fieremans, C., 1991, The kimberlites of Central Africa: a review, in Kampunzu, A. B. and Lubala, R.T., eds., Magmatism in Extensional Structural Settings: The Phanerozoic African Plate: Berlin, Springer, p. 537-559.

François, A., 1974, Stratigraphie, tectonique et minéralisations dans I'Arc cuprifère du Shaba (République du Zaïre), in Bartholomé, P., ed., Gisements stratiformes et provinces cuprifères: Société Géologique de Belgique, p. 79-101.

Höhndorf, A. and Vetter, U., 1999, The Sanyati Ore Deposits in Zimbabwe: Pb - isotopic investigation of sulfide and oxide ores: Zeitschrift für angewande Geologie, v. 45 (1), p. 11-13.

Key, R. M. and Armstrong, R. A., 2000, Geology and geochronology of preKatangan igneous and meta-igneous rocks north of the Lufilian Arc in northwest Zambia: Journal of African Earth Sciences, v. 31, p. 36-37.

Kusky, T. M., 1998, Tectonic setting and terrane accretion of the Archean Zimbabwe craton: Geology, v. 26, p. 163-166.

Lavreau, J. and Deblond, A., 2000, Geological and chronological setting of the greenstone belts of the northern Congo Craton: Journal of African Earth Sciences, v. 30, p. 53.

Lefebvre, J. J., 1975, Les roches ignées dans le Katangien du Shaba (Zaïre). Le district du cuivre: Annales de la Société Géologique de Belgique, v. 98, p. 47-73.

Liyungu, A .K. and Vinyu, M. L., 1996, Constraints on the timing of the high grade Lutembe quartz-feldspathic granulite, charnockitic enderbite and the relationship to the Chipata granite in the Mozambique Belt, in Kamona, A. F., Tembo, F., Mapani, B. S. E. eds., Abstracts Volume, First International Field Conference of IGCP 363 Palaeoproterozoic of SubEquatorial Africa, 14-30 September, 1996, Zambia-Zimbabwe. Lusaka, Geol. Soc. Zambia, p. 19.

Ludwig, K. R., 2000. Users Manual for Isoplot/Ex version 2.3, a geochronological toolkit for Microsof Excel. Berkeley Geochronology Center, Special publication No. 1a.

Master, S., 1991, Stratigraphy, tectonic setting, and mineralization of the early Proterozoic Magondi Supergroup, Zimbabwe: a review: Economic Geology Research Unit Information Circular, Department of Geology, University of the Witwatersrand, Johannesburg, No. 238, 75 pp.

Pinna, P., Cocherie, A., Thieblemont, D., Jezequell, P.and Kayagoma, E., 1999, The Archean evolution of the Tanzanian craton (2.93-2.53 Ga): Journal of African Earth Sciences, v. 28, p. 62-63.

Rainaud, C., Armstrong, R. A., Master, S. and Robb, L. J., 1999, A fertile Palaeoproterozoic magmatic arc beneath the Central African Copperbelt, Mineral deposits: processes to processing: London, in: Stanley, C.J. et al. eds. Mineral Deposits: Processes to Processing, Volume 2: A. A. Balkema, Rotterdam, p. 1427-1430.

Schidlowski, M.and Todt, W., 1998, The Proterozoic Lomagundi carbonate province as a paragon of a 13C- enriched carbonate facies: Geology, radiometric age and geochemical significance: Chinese Science Bulletin, v. 43 (supplement), p. 114.

Tack, L. F. -A., M., Wingate, M. and Deblond, A., 1999, Critical assessment of recent unpublished data supporting a single and united geodynamic evolution of the Sao Francisco-Congo-Tanzania cratonic blocks in the Rodinia configuration: Journal of African Earth Sciences, v. 28, p. 75-76.


[^0]:    * This chapter appeared in Journal of the Geological Society, London, Vol. 160, 2003, pp. 11-14

[^1]:    
    
    
    
    
    
    
    
    
    
    
    
    
    
    
    
     $\begin{array}{llllllllllllllllllllllllllll}0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0\end{array}$
    
     Uncertainties given at the one olevel. $\mathrm{f}_{206} \%$ denotes the percentage of ${ }^{206} \mathrm{~Pb}$ that is common Pb.
    3. Correction for common Pbmade using the measured ${ }^{204} \mathrm{~Pb} / /^{206} \mathrm{~Pb}$ ratio
    
    
    $\qquad$
    

    $$
    \text { For \% Conc., } 100 \% \text { denotes a concordant analysis }
    $$

[^2]:    Notes: 1. Uncertainties given at the one $\sigma$ level.
    2. $\mathrm{f}_{206} \%$ denotes the percentage of ${ }^{206} \mathrm{~Pb}$ that is common Pb .
    3. Correction for common Pb made using the measured ${ }^{204} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ratio
    4. For \% Conc., $100 \%$ denotes a concordant analysis.

[^3]:    Uncertainties given at the one s level.
    $\mathrm{f}_{206} \%$ denotes the percentage of ${ }^{206} \mathrm{~Pb}$ that is common Pb .
    3. Correction for common Pb made using the measured ${ }^{204} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ratio.
    4. For \% Conc., $100 \%$ denotes a concordant analysis.

    Notes:

[^4]:    1. Uncertainties given at the one s level.
    2. $f_{206} \%$ denotes the percentage of ${ }^{206} \mathrm{~Pb}$ that is common Pb .
    3. Correction for common Pb made using the measured ${ }^{204} \mathrm{~Pb} /^{206} \mathrm{~Pb}$ ratio

    Notes

